Using Old Geomagnetic Data to Say Something about the Sun

Leif Svalgaard Oct. 16, 2013

The Russian Observatories



Helsinki [Finland] established on 28th March 1838 by imperial decree of Czar Nicholas I. First director J.J. Nervandor (1805-1848).

Ready for business: July 1st, 1844, continued until 1905. Last few years unreliable because of disturbances from electric tramway traffic started in 1901.

Variometer Invented by Gauss



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The Raw Data

1844-1856 10 minute cadence1856-1897 1 hour cadence

2.5 million observations

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Staff: Director and 12 observers [students]

Many hundred notebooks

Detail of Raw Data

Deklination År 1949 Månad December Timme 11										
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Digitized by pupils in elementary schools all across Finland



Sample Data

Regular daily variation





Magnetic storm

Fig. 10b. The most violent magnetic storm in 1844-1853 occurred on December 20, 1847. The figure depicts declination variations in December 17-22, 1847. The greatest peak-to-peak values of D were about 3° in a few hours being roughly ten times larger than the regular variation shown in Fig. 10a.

Relation to HMF Strength B





Wolf's Discovery: $rD = a + b R_W$



 $dY = H \cos(D) dD$ For small dD

A current system in the ionosphere is created and maintained by solar FUV radiation

The Diurnal Variation of the Declination for Low, Medium, and High Solar Activity





All Geomagnetic Observatories Show the Relationship with Sunspot Number



H Scale Value Problem at Helsinki





Also for Declination in 1886



Figure 4: Diurnal variation of H [left] and of D [right] at Helsinki for three years with sunspot number ~25.

Creates Problems for Assessment of Long-Term Variation of Activity



Bartels' u-index

24-hour running means of the Horizontal Component of the low- & midlatitude geomagnetic field remove most of local time effects and leaves a Global imprint of the Ring Current [Van Allen Belts]:



A quantitative measure of the effect can be formed as a series of the unsigned differences between consecutive days: The InterDiurnal Variability, IDV-index

IDV is strongly correlated with HMF B, but is blind to solar wind speed V







Since we can also estimate solar wind speed from geomagnetic indices [Svalgaard & Cliver, JGR 2007] we can calculate the radial magnetic flux from the total B using the Parker Spiral formula:



There seems to be both a Floor and a Ceiling and most importantly no longterm trend since the 1830s.

Re-evaluation of Cosmic Ray Data

Still problem with the 1880-1890s and generally with low values



Conclusion

- Important to correct data for scale value errors
- Signs of consensus on Cosmic Ray data
- No Modern Grand Maximum
- This is still controversial, hence my current interest