The Mean Field of the Sun

Leif Svalgaard Stanford University Sept. 2, 2011

Crimean Astrophysical Observatory [since 1968]

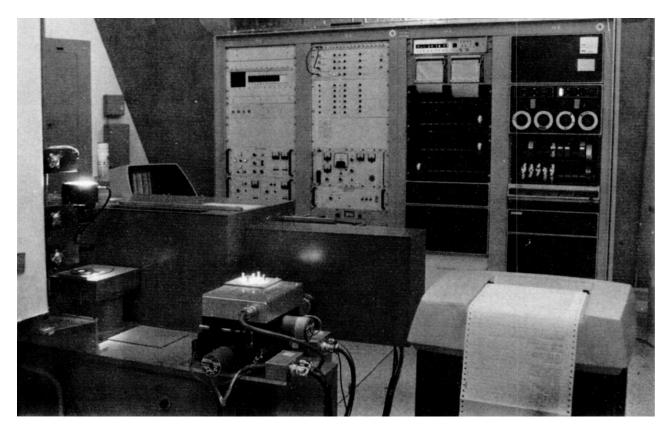
Mount Wilson Observatory [1970-1982]

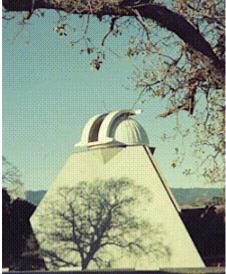






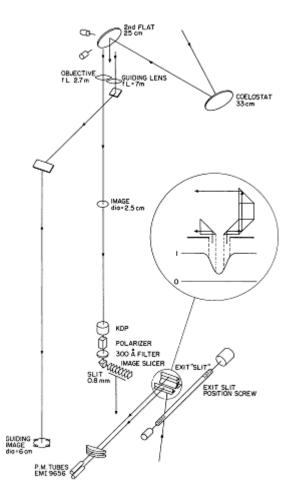
Wilcox Solar Observatory [since 1976]

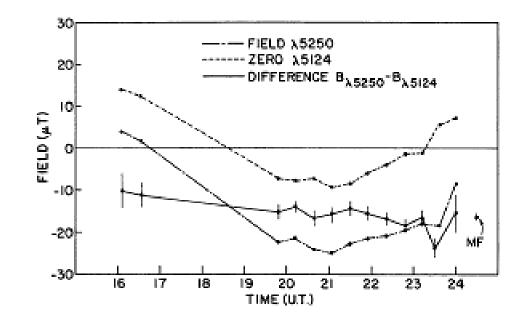




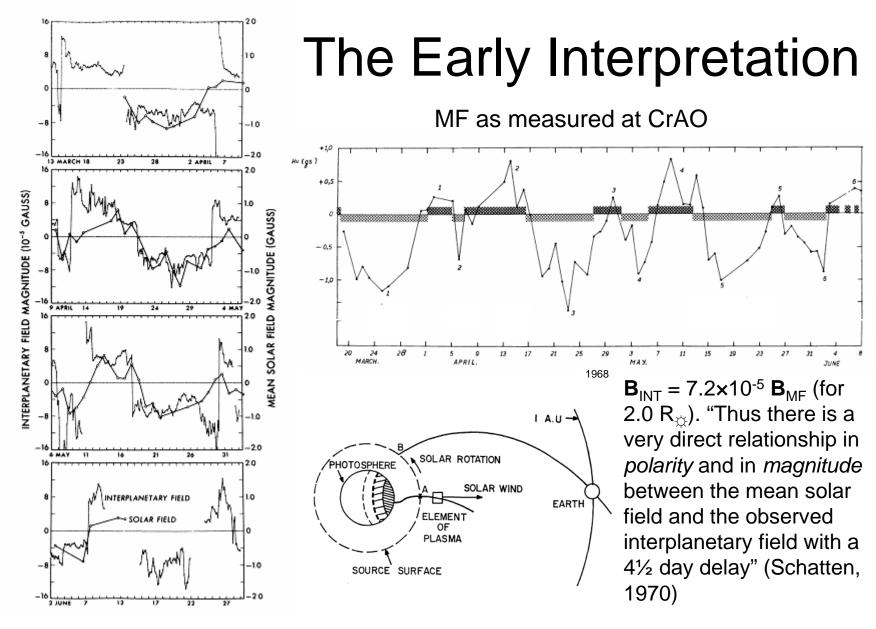


How the Mean Field is measured

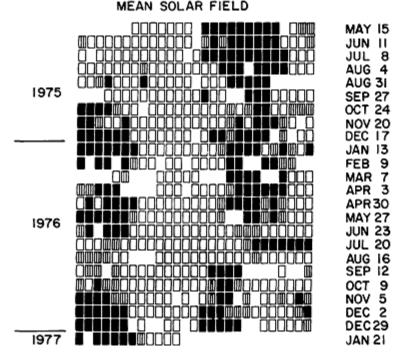




The MF is taken to be the difference between the magnetic signal in λ 525.0 nm and the non-magnetic line λ 512.4 nm.

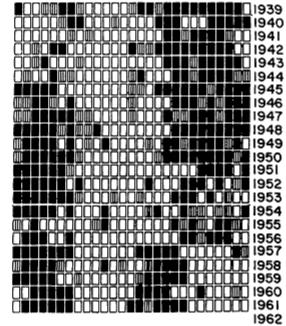


The Polarity Relationship [then]



WSO 5 days before

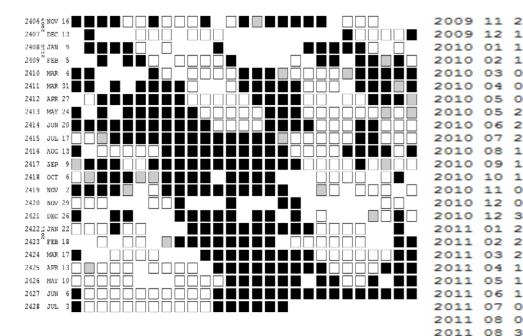
IMF POLARITY



The Polarity Relationship [now]

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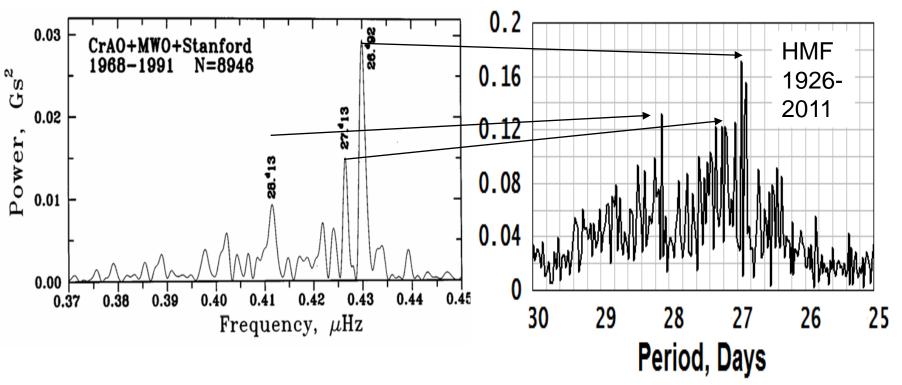


2406	XXX**XXXX*XX.XXXX
2407	*XX
2408	XXXX*XXXXXXXXXXXX
2409	XXXXXX
2410	XXXXXX
2411	XXX*X*XX*X*X.*XX
2412	*XXXXXXXX*XXXXX
2413	XXXXXXXXXX XX*
2414	*XXXXXXXXXXXXX*XX*XX.
2415	.xxxxxxxxxxxxxxx*xxx
2416	*XXXXXXXXXXXXXXXXX*XXX*
2417	.*XXX*X*XXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
2418	*XXXXXXXXXXXXXXXXXXXX
2419	**XXXXXXXXX*XXXX**
2420	X*XXXXXXXXXXX*XXX
2421	X*XXXXXXXXXX****XX
2422	XX*XXX*XXXXXXXXXXX*.X*X
2423	xx.*xxxxxxxxx*xx
2424	**XXX*XXXXX**
2425	*
2426	xxxxx
2427	*XXXXXX*XXXXXXXXXXXXXX
2428	xxxxxxxxxxxxxx*x*
2429	X***XXXXXXXXXXXXXXXXXXXXXX
2430	XX
	2408 2409 2410 2411 2412 2413 2414 2415 2414 2415 2416 2417 2418 2420 2421 2422 2423 2424 2425 2426 2427 2428 2429

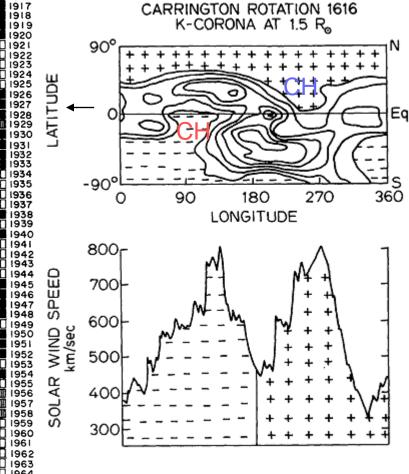
WSO 5 days before

IMF

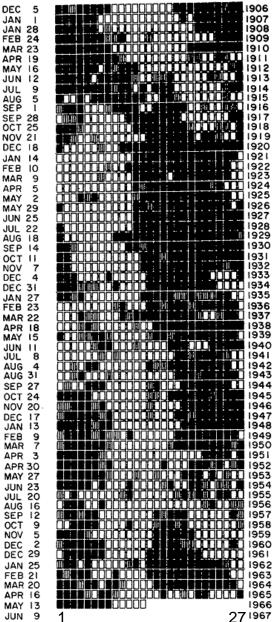
Same Recurrence Periods over Time for MF and HMF [IMF]



Rotation Plots of	
the Sector Polarity	



Skylab Workshop, 1976



IMF POLARITY

Rot - 1st

No

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43 A31

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16 745542 35 64 16

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65 , 65 ,34 ,47

52 A30 27 \$54

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128 654 44 4 22 4 46 5 2 2 2 . 2 24 23 2 777 766 755

A 19 777 761 166 675 4 22 2.4 443 ... 2 576 665 567

12 M16 665 557 54 ... 5 553 2 ... 66 655 525 13 J12 655 325 665 ... 6... 57 652 2... 7 2..... 2

15 45 222 2... 2 2 ... 2.2 576 566 542 ... 342

18 025 . 5765 . . 54 . 62 2 5 . . 675

19 N21 5 ., 25 66 5 16665

19 11 14 14 16, 14. 7665 555 431 1113. 466 53

FID 166 52 13 15 176 666 655 5 15 534 666 424

M2 566 6.4.33 ... 366 655 555 664 343 346 655

26 M29 , 46 655 , 466 556 44, 35, 111 177 654

27 125 177 656 131 678 166 515 353 11. 31 177 556

28 122 ,77 55653 , ... 5 655 554 443 ,67 666

29 418 267 666 62, 556 256 625 542 1.77 56

1930 SK 177 56 676 466 664 456 655 55 . 365 47 67

31 011 . 17 677 776 724 . 65 666 522 2 256 27

32 N7 256 277 66 2 66 2 55 5 66 56 2 . . . 44 2 . . 247

19 031 4, , 6 06, 6 156 522 . 22 3 . . 47. , 16

75 123 645 1. 3 522 165 1776 665 644 24. . . 34 246

1937 M22 . 24 . 26 65 . 2 . . . 2 . 466 766 665 6 546

38 418 6546422 ... 4 54665445 .4 64245

- 6 -

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· 66 6523 .25 26222.

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4 354234. 1362 1 1

65 A16 4 76, 15 5. 474 513 255 4551

1. 2.2 221

127 43. 146 523 616 136 666 656 443 122 645 2.3

66 265 232 .45 3.3 . 5623 22. . 164 224

664 334 333 · 2 · · · · 653 · · · · · 5 664 5.2

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17 528 . . . 534 6 . . 665 556 5 . .

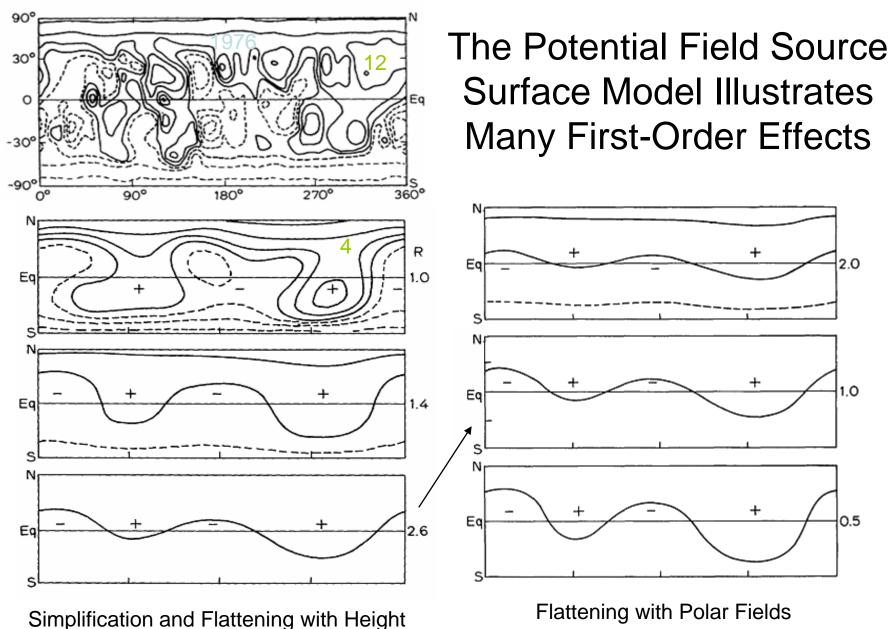
11 11 11 11 17 74 M9 666 414 7 . . 5 767 665 556 554 466 565 446

15 642 111 1924 45 565 446 5 76 766 655 555 541 566 6.4

73 F24 766 755 64 ... 61 222 24 ... 2 57 777 777 665

174 565 427 1910 M23 777 665 446 \$76 ... , 5 77, 777 764

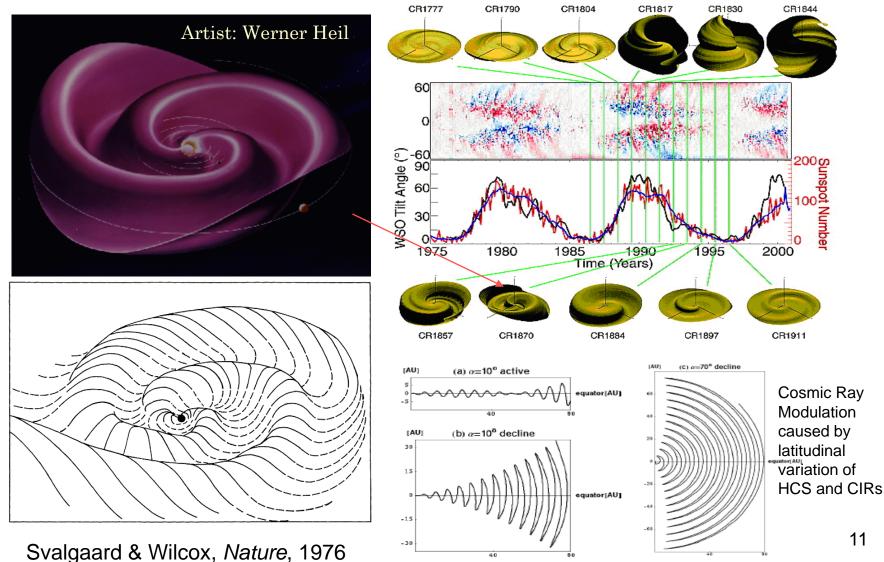
Bartels Rotations



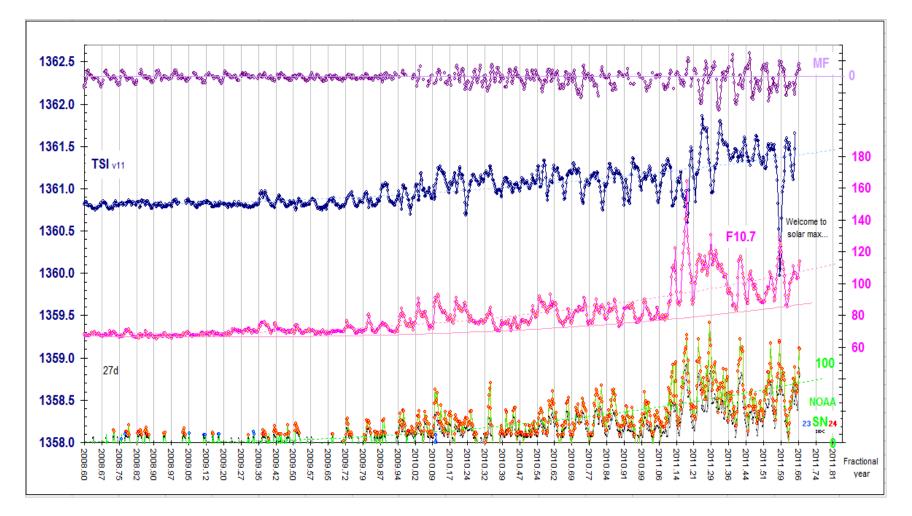
"Domes of closed field lines"

10

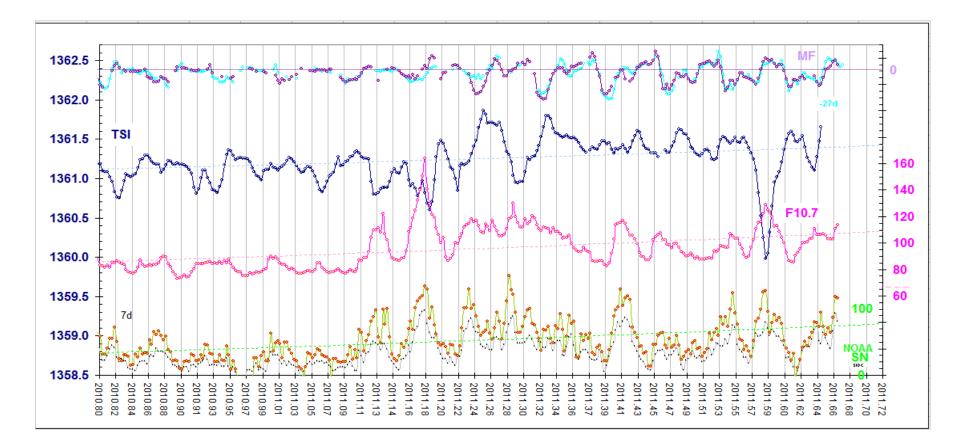
The Heliospheric Current Sheet through the Solar cycle



The MF at the Start of Cycle 24

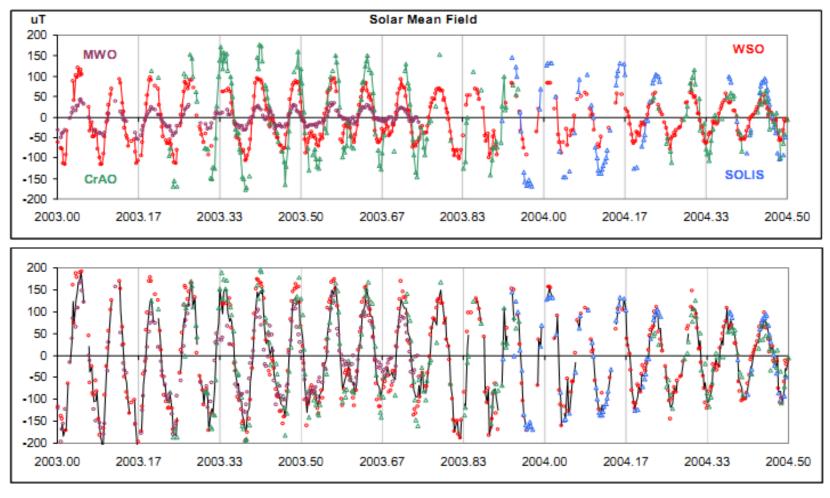


And in Detail



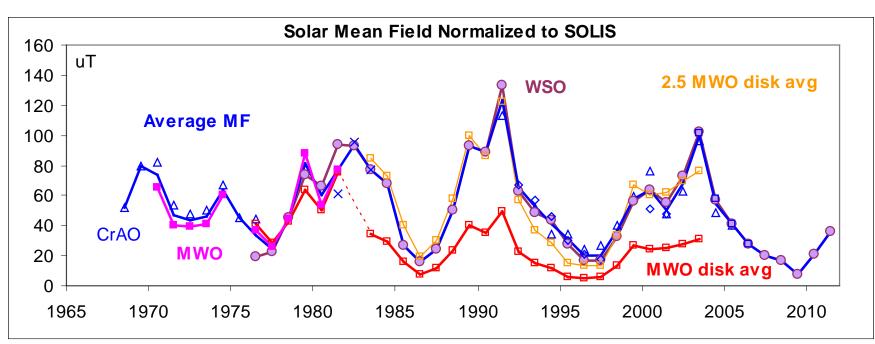
Note the detailed 27-day recurrence in sign and magnitude of the MF

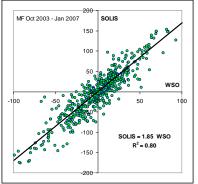
Different Observatories Agree on the Polarity, but NOT on the Magnitude



Roughly: WSO = 2, CrAO = 1, MWO = 4, and SOLIS = 1

MWO Anomaly

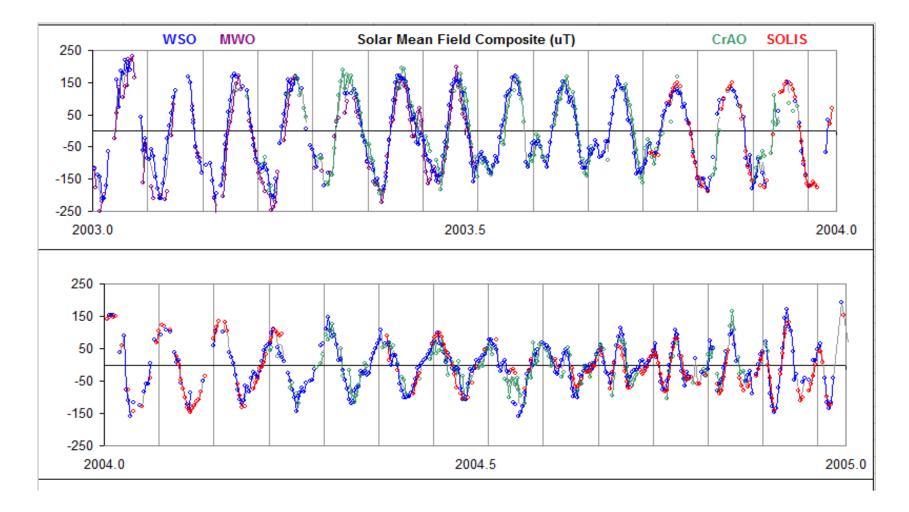




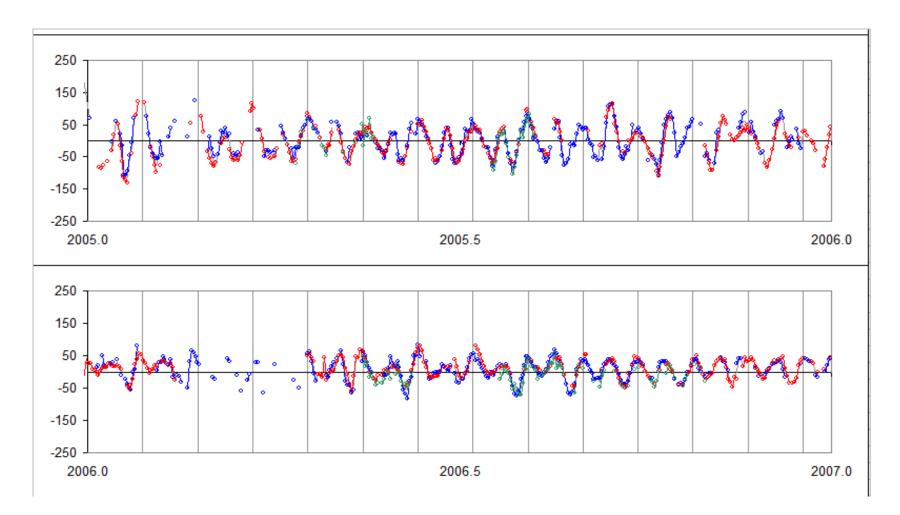
Using the regression factors for each observatory we can bring them all onto the same scale and compute the yearly average of the magnitude

After the upgrade of MWO their MF is much too small

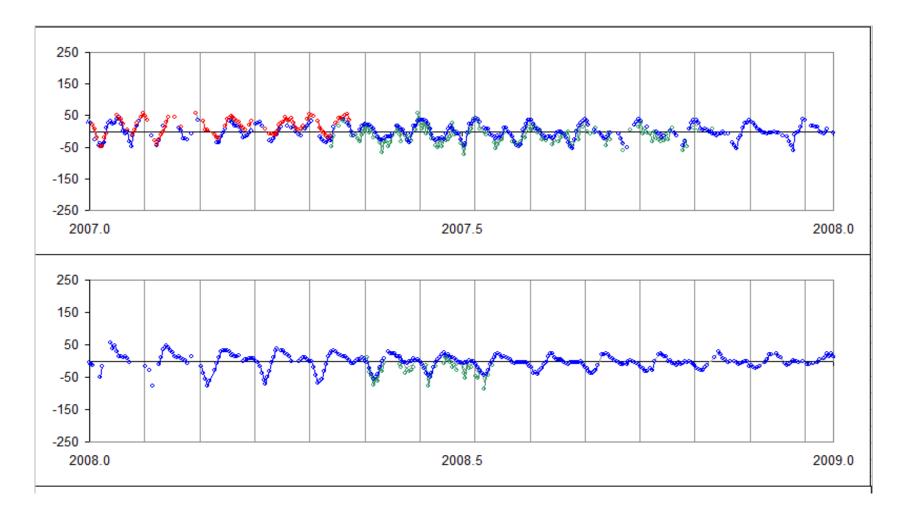
Evolution of the MF Since 2003



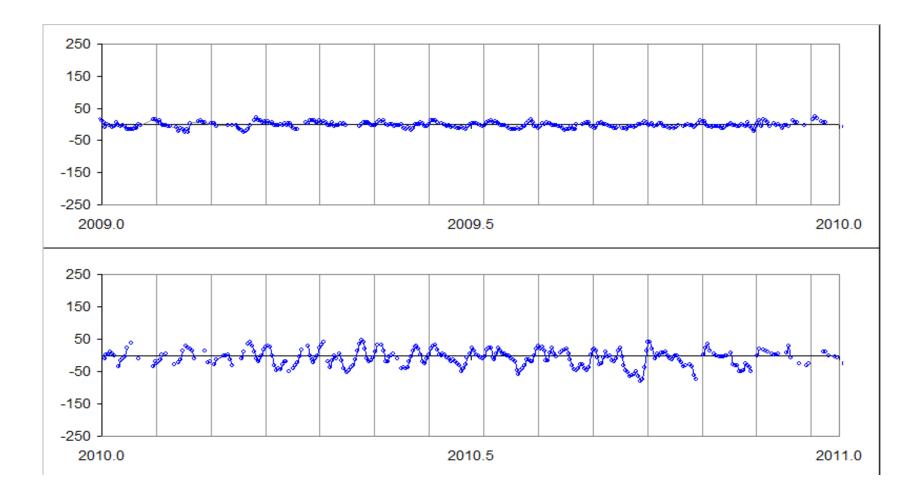
Getting Smaller



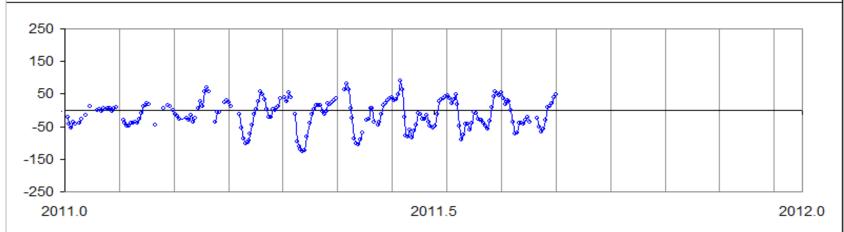
And Smaller

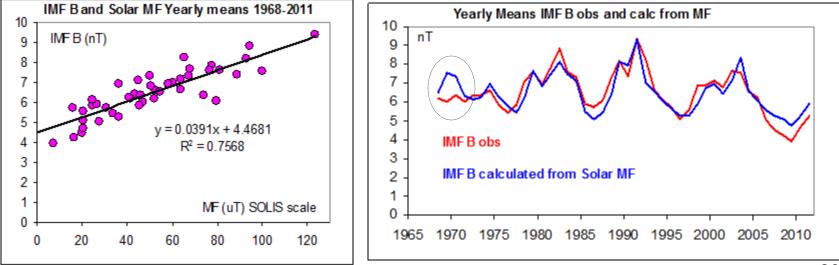


And Smaller, but still matching HMF polarity

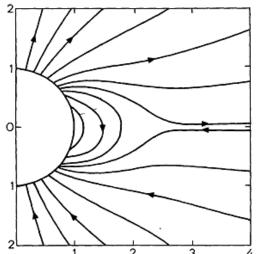


The MF is Riding on a Background HMF that does not Fall Below ~4 nT



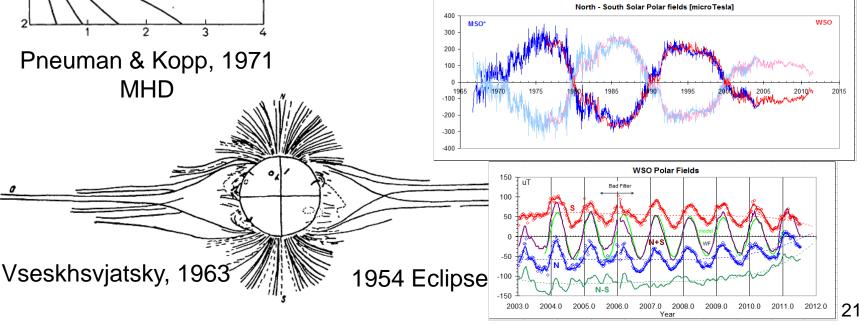


The Importance of the Polar Fields [?]



Even with all the sophistication of current models of the Corona and HMF they are hostage to the correct value of the solar polar fields, which may be different at the two poles and even have longitudinal structure within the polar caps.

This is particularly important at solar minimum when the HCS is largely flat.



Conclusion

- The Solar Mean Field continues to track the polarity of the HMF
- The magnitude of the MF does not reflect that of the HMF, but rides on top of a fixed [?] background [i.e. that does not track the polar fields]
- The MF can be used to monitor the calibration of magnetographs
- We should calculate the MF from HMI as a product.